

AD-A202 080

DTIC FILE COPY

④

OFFICE OF NAVAL RESEARCH

Contract N00014-80-K-0852

R&T Code \_\_\_\_\_

Technical Report No. 47

Utilization of a Highly Correlated Cluster Model for  
Interpretation of Electronic Spectroscopic Data for  
The High-Temperature Superconductors

By

D. E. Ramaker

Prepared for Publication

in the

AIP/AVS Topical Conference Proceedings

George Washington University  
Department of Chemistry  
Washington, D.C. 20052

December, 1988

Reproduction in whole or in part is permitted for  
any purpose of the United States Government

This document has been approved for public release  
and sale; its distribution is unlimited.

DTIC  
ELECTE  
S DEC 19 1988 D  
H

ADA202080

SECURITY CLASSIFICATION OF THIS PAGE

REPORT DOCUMENTATION PAGE			
1a. REPORT SECURITY CLASSIFICATION Unclassified		1b. RESTRICTIVE MARKINGS	
2a. SECURITY CLASSIFICATION AUTHORITY		3. DISTRIBUTION / AVAILABILITY OF REPORT Approved for Public Release, distribution Unlimited.	
2b. DECLASSIFICATION / DOWNGRADING SCHEDULE		5. MONITORING ORGANIZATION REPORT NUMBER(S)	
4. PERFORMING ORGANIZATION REPORT NUMBER(S) Technical Report # 47		7a. NAME OF MONITORING ORGANIZATION Office of Naval Research (Code 413)	
6a. NAME OF PERFORMING ORGANIZATION Dept. of Chemistry George Washington Univ.	6b. OFFICE SYMBOL (If applicable)	7b. ADDRESS (City, State, and ZIP Code) Chemistry Program 800 N. Quincy Street Arlington, VA 22217	
6c. ADDRESS (City, State, and ZIP Code) Washington, D.C. 20052	8a. NAME OF FUNDING / SPONSORING ORGANIZATION Office of Naval Research	8b. OFFICE SYMBOL (If applicable)	9. PROCUREMENT INSTRUMENT IDENTIFICATION NUMBER Contract N00014-80-K-0852
8c. ADDRESS (City, State, and ZIP Code) Chemistry Program 800 North QUINCY, Arlington, VA 22217	10. SOURCE OF FUNDING NUMBERS		
	PROGRAM ELEMENT NO. 61153 N	PROJECT NO.	TASK NO. PP 013-08-01
			WORK UNIT ACCESSION NO NR 056-681
11. TITLE (Include Security Classification) Utilization of a Highly Correlated Cluster Model for Interpretation of Electronic Spectroscopic Data for the High Temperature Superconductors. (UNCL.)			
12. PERSONAL AUTHOR(S) D. E. Ramaker			
13a. TYPE OF REPORT Interim Technical	13b. TIME COVERED FROM TO	14. DATE OF REPORT (Year, Month, Day) December 1988	15. PAGE COUNT 7
16. SUPPLEMENTARY NOTATION Prepared for publication in AIP/AVS Topical Conference Proceedings			
17. COSATI CODES		18. SUBJECT TERMS (Continue on reverse if necessary and identify by block number)	
FIELD	GROUP	SUB-GROUP	
		Superconductivity, Hubbard Model, Photoelectron Spectroscopy, Auger Spectroscopy, Copper Oxides (Mn)	
19. ABSTRACT (Continue on reverse if necessary and identify by block number) We have consistently interpreted electron spectroscopic data for the high temperature superconductors utilizing a highly-correlated $\text{CuO}_n$ cluster model, and an extended Hubbard Hamiltonian which includes the inter-site Cu-O and O-O $U_{pp}$ parameters. The data indicate much larger $U_p$ and $U_{pp}$ values than found in other typical highly conductive metals. Previously unassigned features in the data are now assigned within the model.			
20. DISTRIBUTION / AVAILABILITY OF ABSTRACT <input checked="" type="checkbox"/> UNCLASSIFIED/UNLIMITED <input checked="" type="checkbox"/> SAME AS RPT. <input type="checkbox"/> OTC USERS		21. ABSTRACT SECURITY CLASSIFICATION Unclassified	
22a. NAME OF RESPONSIBLE INDIVIDUAL Dr. David L. Nelson		22b. TELEPHONE (Include Area Code) (202) 696-4410	22c. OFFICE SYMBOL

DD FORM 1473, 84 MAR

83 APR edition may be used until exhausted.  
All other editions are obsolete.

SECURITY CLASSIFICATION OF THIS PAGE  
Unclassified

①

38 12 13 039

UTILIZATION OF A HIGHLY-CORRELATED  $\text{CuO}_x$  CLUSTER MODEL TO  
INTERPRET ELECTRON SPECTROSCOPIC DATA FOR THE HIGH-  
TEMPERATURE SUPERCONDUCTORS

David E. Ramaker<sup>a</sup>  
Department of Chemistry, George Washington University  
Washington, DC 20052, USA

ABSTRACT

We have consistently interpreted electron spectroscopic data for the high temperature superconductors utilizing a highly-correlated  $\text{CuO}_x$  cluster model, and an extended Hubbard Hamiltonian which includes the inter-site  $\text{Cu-O}$   $U_{pd}$  and  $\text{O-O}$   $U_{pp}$  parameters. The data indicate much larger  $U_p$  and  $U_{pp}$  values than found in other typical highly conductive metals. Previously unassigned features in the data are now assigned within the model.

INTRODUCTION

In this work we summarize results of an interpretation of electron spectroscopic data for the high temperature superconductors. The data interpreted include the valence band (VB), Cu 2p, and O 1s photoelectron data (UPS and XPS), the Cu  $L_{2,3}VV$ , Cu  $L_{2,3}M_{2,3}V$ , and O KVV Auger data, and the O K and Cu  $L_{2,3}$  x-ray absorption and emission (XANES and XES) data. Published data for polycrystalline and single crystal samples of  $\text{La}_{1-x}\text{Ba}_x\text{CuO}_4$  and  $\text{YBa}_2\text{Cu}_3\text{O}_{7-x}$  (herein referred to as the La and 123 HTSC's) are considered along with that for  $\text{CuO}$  and  $\text{Cu}_2\text{O}$ .

The basic electronic structure of the HTSC's can be described with the Anderson Hamiltonian utilized by Sawatzky and coworkers<sup>1,2</sup>. It includes the transfer or hopping integral  $t$ , the Cu and O orbital energies  $\epsilon_d$  and  $\epsilon_p$ , the core polarization energy  $Q_d$ , and the intra-site Coulomb repulsion energies  $U_d$  and  $U_p$  (the latter sometimes are assumed to be zero). This model is most useful when the  $U$ 's are large relative to the band widths<sup>1</sup>, i.e. when correlation effects dominate hybridization effects. A  $\text{CuO}_x$ <sup>(2x-1)-</sup> cluster model, which is also reasonably valid when  $U \gg t$ , simplifies the model further<sup>1</sup>. We utilize an extended Hubbard model by adding the inter-site repulsion energies  $U_{dp}$  and  $U_{pp}$ <sup>3</sup> (i.e. between neighboring Cu-O and O-O atoms). The addition of these interactions is important for understanding many of the features in the data.

RESULTS

Our results for the Hubbard parameters are summarized at the top of Table 1. Other estimates of these Hubbard  $U$  and  $c$  parameters have been reported previously for the HTSC's<sup>2-4</sup>. These were obtained empirically from the Cu 2p XPS and the VB UPS data utilizing the Anderson model. Our optimal extended Hubbard

<sup>a</sup>Supported in part by the Office of Naval Research.

results indicate that  $U_p = 12$  and  $U_{pp} = 4.5$  eV for 123. These are much larger than previously thought for these metallic systems, although  $U_p - U_{pp}$  is in agreement with the best theoretical results above.

An upper estimate of the two-center  $pp^*$  hole-hole repulsion,  $U_{pp}^*$ , can be obtained from the Klopman approximation<sup>7</sup>,

$$U_{ij} = e^2/(r_{ij}^2 + (2e^2/(U_i + U_j))^{1/2}), \quad (1)$$

where  $r_{ij}$  is the interatomic distance and  $U_i$  and  $U_j$  are the corresponding intra-atomic repulsion energies. Equation 1 gives a value for  $U_{pp}^*$  around 4.8 eV assuming  $r_{Cu-O}$  is 2.7 Å. The experimental energies of 9.5 and 5.0 eV for the  $pp^*$  and  $pp^*$  features in 123 suggests that the  $pp^*$  final state energy is 7.2 eV. This gives an empirical estimate for  $U_{pp}^*$  of 4.2 eV, very close to the Klopman theoretical result, which does not include the effects of interatomic screening.

The above result shows that metallic screening of two holes, which are spatially separated on neighboring O atoms, is not very significant. This is in contrast to two Cu-O holes, where Table 1 indicates the optimal  $U_{pp} = 1$  eV, while eq. 1 estimates  $U_{pp}$  at 6.1 eV assuming  $r_{Cu-O}$  is 1.9 Å. This large reduction in  $U_{pp}$  may result from charge transfer into the Cu 4sp levels to screen the Cu-O holes. Although metallic screening, which results from virtual electron-hole (e-p) pair excitations at the Fermi level, is not expected to be large in an insulator such as CuO, screening effects are expected to be much larger in metals, such as the HTSC's. The above results show that  $U_{pp}$  is significantly reduced in both, and  $U_{pp}^*$  remains large in both. The lack of a significant change in the  $U$ 's between CuO and the HTSC's indicates that the DOS at the Fermi level in the HTSC's must be very small.

Table 1 also correlates the calculated energies of the excited states with features in the experimental data. The  $CuO_{(1-x)2}$  cluster has one hole shared between the Cu 3d and O 2p shells in the ground state, which we term the  $v$  (valence) states. The spectroscopic final states reflect multi-hole states, e.g.  $v^2$ ,  $cv$  ( $c$  = core) etc. The  $v$  states, as reflected by the theoretical DOS<sup>8</sup>, have the Cu-O bonding ( $\uparrow_b$ ) and antibonding ( $\uparrow_a$ ) orbitals centered at 4 and 0 eV, respectively, and the nonbonding Cu and O orbitals at 2 eV. The O features each have a width  $2\Gamma = 4$  eV due to the O-O bonding and antibonding character and the Cu-O dispersion. We also define the Cu-O hybridization shift  $\delta_1 = [(\Delta^2 + 4\Gamma^2)^{1/2} - \Delta]/2$ , which is utilized in Table 1 to give the energies. Thus, the ground state of an average  $CuO_x$  cluster is located at 1 eV having the energy  $\epsilon_0 - \delta_1 + \Gamma/2 = \epsilon_0 - \alpha$ , which we use as a reference energy for the excited states. In CuO, the hybridization shift  $\Gamma$  is smaller, and we shall see below that  $\Delta = \epsilon_p - \epsilon_d$  has increased to 1 eV.

Those clusters containing additional charge carrier holes (these exists in doped La, and 123 when  $x > 0.5$ ) actually have two holes per  $CuO_x$  cluster. The average  $v^2$  ground state, which is dominated by the  $pp^*$  configuration, i.e. the charge carrier holes spend most of their time on the O atoms, so we indicate this ground state by the

Codes

Dist	Avail and/or Special	
A-1		



TABLE 1 Summary of hole states revealed in the spectroscopic data, and estimated energies using the following optimal values for the Hubbard parameters in eV:

$\delta_1 = 2$	$c_d = 2$	$U_p = 12, 13$	$U_d = 9.5, 10.2$
$\delta_2 = 0.5, 0.8$	$c_p = 2, 3$	$U_{pp} = 4.5, 4$	$U_{dp} = 1$
$\Gamma = 2$	$U_{pp} = 0.$	$U_{cp} = 2$	$Q_d = 9$
$\alpha = 1, 0.5$	$\beta = 2$	$\Delta = 0, 1.$	$K = 4$

State <sup>b</sup>	Energy expression	Calc. E. eV <sup>c,d</sup>	Exp. E. eV <sup>c</sup>	Remark
<u>G.S. and IPES, v</u>				
<sup>†</sup> a) d	$c_d - \delta_1 \mp \Gamma$	0 $\mp$ 2	-	heavily
<sup>†</sup> b) p	$c_p + \delta_1 \mp \Gamma$	4 $\mp$ 2	-	mixed
<u>UPS and XES, v<sup>2</sup></u>				
1) <sup>*</sup> pp <sup>*</sup>	$c_p + \Delta - \delta_2 + \alpha$	2.5	2.5	heavily
2) <sup>*</sup> dp	$c_p + U_{dp} + \delta_2 + \alpha$	4.5	4.2	mixed
3) pp <sup>0</sup>	$c_p + \Delta + U_{pp} - \Gamma + \alpha$	5.5	5.	
4) pp <sup>0</sup>	$c_p + \Delta + U_{pp} + \Gamma + \alpha$	9.5	9.5	mystery peak
5) d <sup>2</sup>	$c_d + U_d + \alpha$	12.5	12.5	Cu sat.
6) p <sup>2</sup>	$c_p + \Delta + U_p + \alpha$	15	16	
<u>Cu 2p XPS, cv</u>				
d $\rightarrow$ cp	$c_c + \Delta + \alpha$	$c_c + 1$	$E_{1p}$	main
cd	$c_c + Q_d + \alpha$	$c_c + 10$	$E_{1p} + 9.2$	sat.
<u>Cu 2p XPS for NaCuO<sub>2</sub>, pp<sup>*</sup> <math>\rightarrow</math> cv<sup>2</sup></u>				
pp <sup>*</sup> -cpp <sup>*</sup>	$c_c + \delta_2 + \beta$	$c_c + 2.5$	$c_c + 2.2$	main
cpp <sup>0</sup>	$c_c + U_{pp} - \Gamma + \delta_2 + \beta$	$c_c + 4.5$	$c_c + 5$	?
cpp <sup>0</sup>	$c_c + U_{pp} + \Gamma + \delta_2 + \beta$	$c_c + 8.5$	$c_c + 9$	?
cdp	$c_c - \Delta + Q_d + U_{dp} + \delta_2 + \beta$	$c_c + 11.5$	$c_c + 11$	sat.
cp <sup>2</sup>	$c_c + U_p + \delta_2 + \beta$	$c_c + 15.5$	$c_c + 14$	sat.?
<u>O 1s XPS, cv</u>				
d $\rightarrow$ cd	$c_c + \alpha$	$c_c + 1$	$E_{1s}$	main
cp <sup>0</sup>	$c_c + \Delta + \alpha$	$c_c + 1$	$E_{1s}$	main
cp <sup>0</sup>	$c_c + \Delta + U_{cp} + \alpha$	$c_c + 3$	$E_{1s} + 2$ ?	tail
cp	$c_c + \Delta + Q_p + \alpha$	?	?	not obs
pp <sup>*</sup> -cdp <sup>0</sup>	$c_c - \Delta + U_{dp} + \delta_2 + \beta$	$c_c + 3.5$	$E_{1s} + 2$ ?	tail
<u>Cu L<sub>2,3</sub> VV AES, v<sup>2</sup></u>				
dpp <sup>*</sup>	$2c_p + 2U_{dp} + \alpha$	7	7	2 cent.
dpp <sup>0</sup>	$2c_p + U_{pp} + 2U_{dp} + \alpha$	11.5	-	no mix
d <sup>2</sup> p	$c_d + c_p + U_d + 2U_{dp} - \delta_2 + \alpha$	16	15.5	main
dp <sup>2</sup>	$2c_p + U_p + 2U_{dp} + \delta_2 + \alpha$	19.5	18-25	sat.

TABLE 1 (cont.)

State <sup>a</sup>	Energy expression	Calc. E. eV <sup>c,d</sup>	Exp. E. eV <sup>c</sup>	Remark
<b>Cu L<sub>2,3</sub>V AES, cv<sup>1</sup></b>				
cdp	$\epsilon_c + \epsilon_p + Q_d + U_d + \frac{1}{2}K + \alpha$	$\epsilon_c + 9$	$E_{2p} + 10$	main, <sup>1</sup> L
		$\epsilon_c + 17$	$E_{2p} + 18$	main, <sup>1</sup> L
cp <sup>2</sup>	$\epsilon_c + \epsilon_p + \Delta + U_p + \alpha$	$\epsilon_c + 15$	-	not
cd <sup>2</sup>	$\epsilon_c - \epsilon_d + U_d + 2Q_d + \alpha$	$\epsilon_c + 30.5$	-	obs.
<b>Cu L<sub>2,3</sub> EELS, c</b>				
d → c	$\epsilon_c - \epsilon_d + \delta_1$	$E_{2p} - 1$	$E_{2p} - 1.4$	edge
cpCB	$\epsilon_c + \Delta - CB + \alpha$	$E_{2p} - CB$	$E_{2p} + 1.2$	upper
pp <sup>2</sup> → cp	$\epsilon_c - \epsilon_p + \delta_2 + \delta$	$E_{2p} - 0.5$	$E_{2p}$	middle
<b>O K EELS, c</b>				
d → c	$\epsilon_c - \epsilon_d + \delta_1$	$E_{1s} - 1$	$E_{1s}$	edge
cdCB	$\epsilon_c - CB + \alpha$	$E_{1s} - CB$	$E_{1s} + 1.7$	upper
pp <sup>2</sup> → cd	$\epsilon_c - \Delta - \epsilon_p + \delta_2 + \delta$	$E_{1s} - 0.5$	-	not obs

<sup>a</sup>Parameters for 123 indicated first, those for CuO second.

<sup>b</sup>The dominant character in the hybridized states is given.

<sup>c</sup>The Calc. E and Exp. E columns indicate the results for 123, except for the "Cu 2p XPS, pp<sup>2</sup> → cv<sup>1</sup>" section, which is for NaCuO<sub>2</sub>.

<sup>d</sup>The calculated E is defined relative to the ground v<sup>1</sup> (d) state energy =  $\epsilon_d - \alpha$ , or to the v<sup>1</sup> (pp<sup>2</sup>) ground state energy =  $2\epsilon_p - \delta_2 - \delta$ . The v<sup>1</sup>(d) energy defines the Fermi level relative to the vacuum level at zero.

<sup>e</sup>The dominant character switches as described in the text, and thus the sign in front of  $\delta_2$  is the opposite for CuO.

parameters in Table 1 were obtained by considering this same data plus XANES, Auger and XES data. Although we are in general agreement with the reported magnitudes for most of the parameters, our  $U_d$  value is larger by about 2-3 eV so that it is consistent with the AES data. In Table 1, we indicate the location of two valence holes by d (Cu 3d) or p (O 2p). In the case of two holes on the oxygens, we distinguish two holes on the same O (p<sup>2</sup>), on ortho neighboring O atoms (pp<sup>2</sup>), or on para O atoms (pp<sup>2</sup>) of the cluster. Furthermore, neighboring pp<sup>2</sup> holes can dimerize<sup>4</sup>, so we distinguish between two holes in bonded (pp<sup>2</sup><sub>b</sub>) and antibonded (pp<sup>2</sup><sub>a</sub>) O pairs.

The magnitudes of the U parameters are critical to the mechanism for the superconductivity. As a consequence, much effort has also gone into theoretically calculating these parameters, but wide disagreement still exists over the magnitudes. Theoretical values for  $U_d$  in the range 6.5-10 eV,  $U_p$  (actually  $U_p - U_{pp}$ ) in the range 7-14 eV, and  $U_{pp}$  in the range 0.6-1.6 eV have been reported<sup>4</sup>, with the smaller results favored based on the quality of the calculations. No results for  $U_{pp}$  have been reported. Our empirical

notation  $pp^*$ . We use  $2\epsilon_p - \delta_p - \delta$  as the energy of the  $pp^*$  ground state relative to the vacuum level, where  $\delta = 2$  eV is the energy shift between the principal  $pp^*$  UPS final state at 2.5 eV and the lowest ground  $pp^*$  states around 0.5 eV from the Fermi level.

The correlation between the calculated energies and experimental features, utilizing the indicated optimal Hubbard parameters is very good. Details of this work are published elsewhere<sup>9</sup>. Figs. 1 and 2 show examples of the UPS and Cu AES data for 123 and  $\text{CuO}$ , which reveal some of the features itemized in Table 1; the remaining data are published elsewhere.

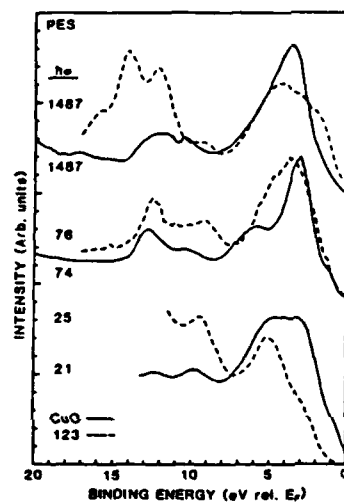


Figure 1. Comparison of UPS spectra for  $\text{CuO}$  and 123 taken with the indicated photon energies in eV. Data for  $\text{CuO}$  from refs. 10 ( $h\nu = 1487$ ), 11 ( $h\nu = 74$ ) and 12 ( $h\nu = 21$ ). Data for 123 from ref. 13 ( $h\nu = 25$  and  $74$ ) and 14 ( $h\nu = 1487$ ).

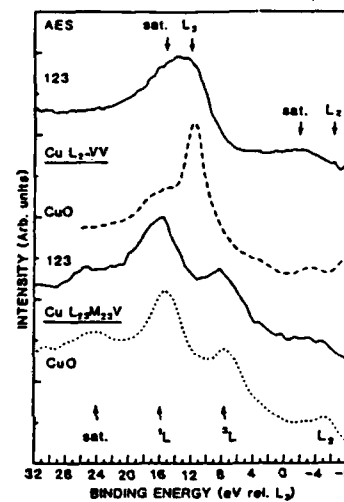


Figure 2. Comparison of Auger data for the materials indicated.  $L_{23}VV$  data for  $\text{CuO}$  and 123 from ref. 15.  $L_{23}M_{23}V$  data for  $\text{CuO}$  from ref. 16 and for 123 from Ref. 9. The  $L_{23}VV$  data is on a 2-hole binding energy scale  $= E_{L_2} - E_{L_1}$ , and the  $L_{23}M_{23}V$  on a 1-hole scale  $= E_{L_2} - E_{L_1} - E_{M_{23}}$ , where  $E_{L_1} = 933.4$  and  $E_{M_{23}} = 77.3$  eV<sub>123</sub>.

#### SUMMARY AND CONCLUSIONS

This work has allowed us to assign some previously unassigned features in the data, and greatly increased our understanding of the

dynamical electronic processes which produce these features. We itemize our conclusions as follows:

1) A switch in the character of state 1 (see Table 1) from more  $dp$  to  $pp^*$  and vice versa for state 2 between  $\text{CuO}$  and 123 arises because  $\Delta$  decreases from 1 eV to 0 eV. The smaller  $\Delta$  in 123, due to a smaller  $\epsilon$ , polarization energy, is consistent with the Cu 2p XPS and XES data (the latter showing this effect dramatically).<sup>17,18</sup> Since state 1 is more of  $pp^*$  character in the SC's, the "charge carrier holes" (present in the La after Sr doping and in the 123 when  $7-x$  is greater than 6.5) spend more time on the oxygens in 123 than in  $\text{CuO}$ .

2) The  $pp^*$  state is believed to be responsible for the "mystery" peak found at 9.5 eV in the UPS. Figure 1 indicates that such a feature also appears for  $\text{CuO}$ .<sup>11,12</sup> This feature does not appear for  $\text{Cu}_2\text{O}$ , as expected since UPS reflects the one-hole DOS in  $\text{Cu}_2\text{O}$ . Thus this feature is not unique to the SC's; it naturally appears for those systems with two-hole photoemission final states.

3) Although  $\text{Cu}_2\text{O}$ ,  $\text{CuO}$ , and  $\text{NaCuO}_2$  have a formal Cu valence of +1, +2, and +3, in the current picture they reflect the  $cv^*$  DOS, with  $n=0$ , 1, and 2. Furthermore, we consistently predict the "chemical shifts" in the primary Cu 2p XPS peaks. Whereas,  $\text{Cu}_2\text{O}$  exhibits just a primary core hole  $c$  state at energy  $\epsilon_c$ ,  $\text{CuO}$  has its primary  $cd$  feature energy shifted by  $\Delta+\alpha$  relative to  $\epsilon_c$ , and the primary  $cpp^*$  feature for  $\text{NaCuO}_2$  by  $\delta_1+\alpha$  (Table 1), which is consistent with the experimental data.<sup>17</sup> The width of the primary feature is seen to correlate with the intensity of the satellite, and is not due to the O p band width as suggested by others.<sup>18</sup>

4) The increased "satellite" feature at 19 eV in the Cu  $L_{2,3}VV$  Auger line shape for the HTSC's compared with  $\text{CuO}$ .<sup>15,16</sup> (see Fig. 2) arises because of increased final-state configuration mixing between the  $d^3p$  and  $dp^1$  states. Its intensity is increased in 123 relative to  $\text{CuO}$  because the energy separation (before hybridization) between  $d^3p$  and  $dp^1$  has decreased from 3.8 eV in  $\text{CuO}$  to 2.5 eV in 123. We have indicated this mixing in Table 1 by adding the hybridization shifts  $\delta_2$  to the energy expressions for these two states.

5) We find that the initial-core shakeup (ICSU) process, which is known to be responsible for the satellite features in the Cu 2p XPS data,<sup>1</sup> does not produce satellites in the AES or XES data, because the ICSU states generally "relax" to the primary states of the same symmetry before the core level decay. Such a relaxation is expected when the ICSU excitation energy is larger than the core level width.<sup>19</sup> Previously, vanderLaan et al.,<sup>1</sup> for the Cu halides, suggested that the intensity of these ICSU states in the XPS should be quantitatively reflected in the intensity of the Auger satellites found in the  $L_{2,3}VV$  lineshapes. The data do not indicate this however. We previously<sup>15</sup> indicated that a fraction of these ICSU states probably resulted in Auger satellites for the HTSC's, and that this fraction becomes larger as the covalency of the HTSC material increases. This work indicates rather that the ICSU states relax before the core level decay to states of the same symmetry, provided they have a ICSU excitation energy that is much greater



than the core level width. We believe this to be a general result, at least in the  $\text{Cu}^{+2}$  materials.

6) The EELS and XANES data<sup>20,21</sup> reflect the contributions from three possible transitions; the dominant  $d \rightarrow c$  contribution nearest the Fermi level, the  $pp^* \rightarrow cv$  ( $v = d$  or  $p$ ) contribution resulting from the carrier hole states, as well as the  $cvCB$  contribution well above the Fermi level<sup>22</sup>. Here  $CB$  represents an electron present in the higher Cu  $4sp$  or O  $3p$  "conduction band". The latter two contributions are not always resolved, and sometimes have been confused in the literature<sup>23-25</sup>.

7) All of the temperature effects seen in the spectroscopic data<sup>22-25</sup> can be attributed to a single phenomenon, namely a decrease in  $\epsilon_p$  due to increased metallic screening, or long range polarization. This is consistent with the decrease in the primary  $cp$  peak energy in the Cu  $2p$  XPS, while the  $cd$  satellite remains unshifted. The larger energy separation between the  $cd$  and  $cp$  states decreases the mixing which causes the satellite to decrease in intensity and the main peak to get narrower. Although the primary  $cd$  peak does not shift in the O  $1s$  XPS, a slight shift to lower energy is seen in the  $cp^*$  and  $cp^*$  contributions at lower temperature, as expected with a decrease in  $\epsilon_p$ . The UPS spectra show a skewing toward the Fermi level at lower temperature, as expected with a decrease in  $\epsilon_p$ . Finally the growth of the satellite intensity in the Cu  $L_{2,3}VV$  Auger lineshape is consistent with a decrease in  $\epsilon_p$ . The increased metallic like screening or polarization which appears to occur at lower temperature, reducing  $\epsilon_p$ , probably involves the grain boundaries, since the more recent data for the single crystal samples do not change with temperature<sup>27</sup>.

In summary, an interpretation of the data utilizing a highly correlated  $\text{CuO}_2$  cluster model shows that a single set of Hubbard parameters predicts all of the state energies. Changes in the data between  $\text{CuO}$  and the HTSC's arises primarily from a reduction in  $\epsilon_p$ ; this reduction continues with decreasing temperature in the HTSC's due to increased metallic screening. Compared with  $\text{CuO}$ , the HTSC's show an increased covalent interaction between the Cu-O bonds. The large size of  $U_{pp^*}$ , and the temperature dependence, reveal that metallic screening is incomplete, and hence that the DOS at the Fermi level in the HTSC's is relatively small.

#### REFERENCES

1. G. vanderLaan et al., Phys. Rev. 24, 4369 (1981); G.A. Sawatzky et al., Phys. Rev. Lett. 53, 2339 (1985); J. Zaanen et al., Phys. Rev. B33, 8060 (1986).
2. Z. Shen et al., Phys. Rev. B36, 8414 (1987).
3. A. Fujimori et al., Phys. Rev. B35, 8814 (1987).
4. J.C. Fuggle et al., Phys. Rev. B37, 1123 (1988).
5. R.A. de Groot, H. Gutfreund, and M. Weger, Sol. State Commun. 63, 451 (1987); W. Folkerts et al., J. Phys. C: Solid State Phys. 20, 4135 (1987); A. Manthiram, X.X. Tang, and J.B. Goodenough, Phys. Rev. B37, 3734 (1988).

6. C.F. Chen et al., unpublished; A.K. McMahan, R.M. Martin, and S. Satpathy, unpublished; M. Schluter, M.S. Hybertsen, and N. E. Christensen, Proc. Intn. Conf. High T<sub>c</sub> Superconductors and Materials and Mechanisms of Superconductivity, J. Muller and J.L. Olsen, Eds., (Interlaken, Switzerland, 1988).
7. G. Klopman, J. Am. Chem. Soc. **86**, 4550 (1964).
8. J. Redinger et al., Phys. Lett. **124**, 463 and 469 (1987).
9. D.E. Ramaker, N.H. Turner, and F.L. Hutson, submitted.
10. A. Rosencwaig and G.K. Wertheim, J. Elect. Spectrosc. Related Phenom. **1**, 493 (1972/73).
11. M.R. Thuler, R.L. Benbow, and Z. Hurych, Phys. Rev. **B26**, 669 (1982).
12. C. Benndorf et al., J. Electron. Spectrosc. Related Phenom. **19**, 77 (1980).
13. N.G. Stoffel et al., Phys. Rev. **B37**, 7952 (1988); **B38**, July (1988).
14. D.C. Miller et al., in Thin Film Processing and Characterization of High Temperature Superconductors, J.M. Harper, J.H. Colton, and L.C. Feldman, Eds., AVS Series No. 3 (AIP: New York, 1988) p 336.
15. D.E. Ramaker et al., Phys. Rev. **36**, 5672 (1987).
16. P.E. Larson, J. Electron Spectrosc. Related Phenom. **4**, 213 (1974).
17. P. Steiner et al., Z. Phys. B- Condensed Matter **67**, 497 (1987).
18. D.D. Sarma, Phys. Rev. **B37**, 7948 (1988).
19. J.W. Gadzuk and M. Sunjic, Phys. Rev. **B12**, 524 (1975).
20. D.D. Sarma et al., Phys. Rev. **B37**, 9784 (1988).
21. N. Nucker et al., Z. Phys. B: Cond. Matter **67**, 9 (1987); Phys. Rev. **37**, 5158 (1988).
22. A. Bianconi et al., Solid State Commun. **63**, 1009 (1987); Intn. J. Modern Phys. **131**, 853 (1987).
23. N.S. Kohiki and T. Hamada, Phys. Rev. **B36**, 2290 (1987).
24. B. Dauth et al., Z. Phys. B- Condensed Matter **68**, 407 (1987).
25. D.H. Kim et al., Phys. Rev. **B37**, 9745 (1988).
26. A. Balzarotti et al., Phys. Rev. **B36**, 8285 (1987).
27. J. Weaver and P. Steiner, private communication.

DL/1113/87/2

TECHNICAL REPORT DISTRIBUTION LIST, GEN

	<u>No. Copies</u>		<u>No. Copies</u>
Office of Naval Research Attn: Code 1113 800 N. Quincy Street Arlington, Virginia 22217-5000	2	Dr. David Young Code 334 NORDA NSTL, Mississippi 39529	1
Dr. Bernard Dorda Naval Weapons Support Center Code 50C Crane, Indiana 47522-5050	1	Naval Weapons Center Attn: Dr. Ron Atkins Chemistry Division China Lake, California 93555	1
Naval Civil Engineering Laboratory Attn: Dr. R. W. Drisko, Code L52 Port Hueneme, California 93401	1	Scientific Advisor Commandant of the Marine Corps Code RD-1 Washington, D.C. 20380	1
Defense Technical Information Center Building 5, Cameron Station Alexandria, Virginia 22314	12 high quality	U.S. Army Research Office Attn: CRD-AA-IP P.O. Box 12211 Research Triangle Park, NC 27709	1
DTNSRDC Attn: Dr. H. Singerman Applied Chemistry Division Annapolis, Maryland 21401	1	Mr. John Boyle Materials Branch Naval Ship Engineering Center Philadelphia, Pennsylvania 19112	1
Dr. William Tolles Superintendent Chemistry Division, Code 6100 Naval Research Laboratory Washington, D.C. 20375-5000	1	Naval Ocean Systems Center Attn: Dr. S. Yamamoto Marine Sciences Division San Diego, California 91232	1

## ABSTRACTS DISTRIBUTION LIST, 056/625/629

Dr. F. Carter  
Code 6170  
Naval Research Laboratory  
Washington, D.C. 20375-5000

Dr. Richard Colton  
Code 6170  
Naval Research Laboratory  
Washington, D.C. 20375-5000

Dr. Dan Pierce  
National Bureau of Standards  
Optical Physics Division  
Washington, D.C. 20234

Dr. R. G. Wallis  
Department of Physics  
University of California  
Irvine, California 92664

Dr. D. Bamaker  
Chemistry Department  
George Washington University  
Washington, D.C. 20052

Dr. J. C. Hemminger  
Chemistry Department  
University of California  
Irvine, California 92717

Dr. T. F. George  
Chemistry Department  
University of Rochester  
Rochester, New York 14627

Dr. G. Rubloff  
IBM  
Thomas J. Watson Research Center  
P.O. Box 218  
Yorktown Heights, New York 10598

Dr. J. Baldeschwieler  
Department of Chemistry and  
Chemical Engineering  
California Institute of Technology  
Pasadena, California 91125

Dr. Galen D. Stucky  
Chemistry Department  
University of California  
Santa Barbara, CA 93106

Dr. A. Steckl  
Department of Electrical and  
Systems Engineering  
Rensselaer Polytechnic Institute  
Troy, New York 12181

Dr. John T. Yates  
Department of Chemistry  
University of Pittsburgh  
Pittsburgh, Pennsylvania 15260

Dr. R. Stanley Williams  
Department of Chemistry  
University of California  
Los Angeles, California 90024

Dr. R. P. Messmer  
Materials Characterization Lab.  
General Electric Company  
Schenectady, New York 22217

Dr. J. T. Keiser  
Department of Chemistry  
University of Richmond  
Richmond, Virginia 23173

Dr. R. W. Plummer  
Department of Physics  
University of Pennsylvania  
Philadelphia, Pennsylvania 19104

Dr. E. Yeager  
Department of Chemistry  
Case Western Reserve University  
Cleveland, Ohio 44106

Dr. M. Winograd  
Department of Chemistry  
Pennsylvania State University  
University Park, Pennsylvania 16802

Dr. Roald Hoffmann  
Department of Chemistry  
Cornell University  
Ithaca, New York 14853

Dr. Robert L. Whetten  
Department of Chemistry  
University of California  
Los Angeles, CA 90024

Dr. Daniel M. Neumark  
Department of Chemistry  
University of California  
Berkeley, CA 94720

Dr. G. H. Morrison  
Department of Chemistry  
Cornell University  
Ithaca, New York 14853

ABSTRACTS DISTRIBUTION LIST, 056/625/629

Dr. J. E. Jensen  
Hughes Research Laboratory  
3011 Malibu Canyon Road -  
Malibu, California 90265

Dr. J. H. Weaver  
Department of Chemical Engineering  
and Materials Science  
University of Minnesota  
Minneapolis, Minnesota 55455

Dr. A. Reisman  
Microelectronics Center of North Carolina  
Research Triangle Park, North Carolina  
27709

Dr. M. Grunze  
Laboratory for Surface Science  
and Technology  
University of Maine  
Orono, Maine 04469

Dr. J. Butler  
Naval Research Laboratory  
Code 6115  
Washington D.C. 20375-5000

Dr. L. Interante  
Chemistry Department  
Rensselaer Polytechnic Institute  
Troy, New York 12181

Dr. Irvin Heard  
Chemistry and Physics Department  
Lincoln University  
Lincoln University, Pennsylvania 19352

Dr. K. J. Klaubunde  
Department of Chemistry  
Kansas State University  
Manhattan, Kansas 66506

Dr. C. B. Harris  
Department of Chemistry  
University of California  
Berkeley, California 94720

Dr. R. Bruce King  
Department of Chemistry  
University of Georgia  
Athens, Georgia 30602

Dr. R. Reeves  
Chemistry Department  
Rensselaer Polytechnic Institute  
Troy, New York 12181

Dr. Steven M. George  
Stanford University  
Department of Chemistry  
Stanford, CA 94305

Dr. Mark Johnson  
Yale University  
Department of Chemistry  
New Haven, CT 06511-8118

Dr. W. Knauer  
Hughes Research Laboratory  
3011 Malibu Canyon Road  
Malibu, California 90265

Dr. Theodore E. Madey  
Surface Chemistry Section  
Department of Commerce  
National Bureau of Standards  
Washington, D.C. 20234

Dr. J. E. Demuth  
IBM Corporation  
Thomas J. Watson Research Center  
P.O. Box 218  
Yorktown Heights, New York 10598

Dr. M. G. Lagally  
Department of Metallurgical  
and Mining Engineering  
University of Wisconsin  
Madison, Wisconsin 53706

Dr. R. P. Van Duyne  
Chemistry Department  
Northwestern University  
Evanston, Illinois 60637

Dr. J. M. White  
Department of Chemistry  
University of Texas  
Austin, Texas 78712

Dr. Richard J. Saykally  
Department of Chemistry  
University of California  
Berkeley, California 94720

ABSTRACTS DISTRIBUTION LIST, 056/625/629

Dr. G. A. Somorjai  
Department of Chemistry  
University of California  
Berkeley, California 94720

Dr. J. Murday  
Naval Research Laboratory  
Code 6170  
Washington, D.C. 20375-5000

Dr. W. T. Peria  
Electrical Engineering Department  
University of Minnesota  
Minneapolis, Minnesota 55455

Dr. Keith H. Johnson  
Department of Metallurgy and  
Materials Science  
Massachusetts Institute of Technology  
Cambridge, Massachusetts 02139

Dr. S. Sibener  
Department of Chemistry  
James Franck Institute  
5640 Ellis Avenue  
Chicago, Illinois 60637

Dr. Arold Green  
Quantum Surface Dynamics Branch  
Code 3817  
Naval Weapons Center  
China Lake, California 93555

Dr. A. Wold  
Department of Chemistry  
Brown University  
Providence, Rhode Island 02912

Dr. S. L. Bernasek  
Department of Chemistry  
Princeton University  
Princeton, New Jersey 08544

Dr. W. Kohn  
Department of Physics  
University of California, San Diego  
La Jolla, California 92037

Dr. Stephen D. Kevan  
Physics Department  
University Of Oregon  
Eugene, Oregon 97403

Dr. David M. Walba  
Department of Chemistry  
University of Colorado  
Boulder, CO 80309-0215

Dr. L. Kasmodel  
Department of Physics  
Indiana University  
Bloomington, Indiana 47403

Dr. K. C. Janda  
University of Pittsburg  
Chemistry Building  
Pittsburg, PA 15260

Dr. E. A. Irene  
Department of Chemistry  
University of North Carolina  
Chapel Hill, North Carolina 27514

Dr. Adam Heller  
Bell Laboratories  
Murray Hill, New Jersey 07974

Dr. Martin Fleischmann  
Department of Chemistry  
University of Southampton  
Southampton SO9 5NH  
UNITED KINGDOM

Dr. H. Tachikawa  
Chemistry Department  
Jackson State University  
Jackson, Mississippi 39217

Dr. John W. Wilkins  
Cornell University  
Laboratory of Atomic and  
Solid State Physics  
Ithaca, New York 14853

Dr. Ronald Lee  
R301  
Naval Surface Weapons Center  
White Oak  
Silver Spring, Maryland 20910

Dr. Robert Gomer  
Department of Chemistry  
James Franck Institute  
5640 Ellis Avenue  
Chicago, Illinois 60637

Dr. Noria Metiu  
Chemistry Department  
University of California  
Santa Barbara, California 93106

Dr. W. Goddard  
Department of Chemistry and Chemical  
Engineering  
California Institute of Technology  
Pasadena, California 91125